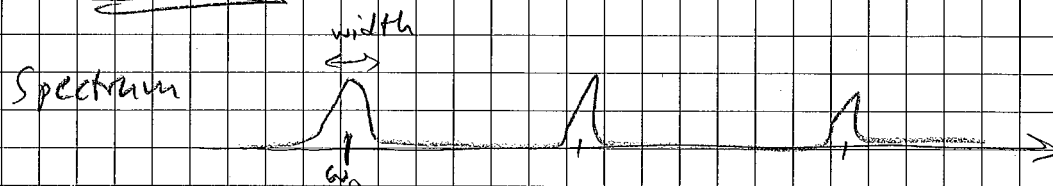
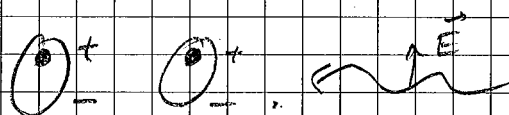


2.5-2.7

(4)

Broadening & Bloch equations

4-1

Atoms absorb & emit lighteven if it is not resonant- if we take A_2 into account! $\Rightarrow$ linewidthHow?An EM-wave will polarize the atomDef. Polarization $\vec{P}(t) = \frac{N}{V} (-e \vec{D}(t))$ Assume $\vec{E} \parallel \vec{e}_x$, then the magnitude

$$P(t) = -e \frac{N}{V} X(t) \quad [\text{Recall } X = \text{c.m.}]$$

Write

$$P(t) = \frac{1}{2} \epsilon_0 E_0 \{ X(\omega) e^{i\omega t} + X(-\omega) e^{-i\omega t} \}$$

where we definedsusceptibility

$$X(\omega) = \underbrace{X'(\omega)}_{=\text{Re } X} + i \underbrace{X''(\omega)}_{=\text{Im } X}$$

χ' real part $\leftrightarrow$ refraction

4-2

χ'' im part $\leftrightarrow$ absorption

Solve Schrödinger eq.

Include spontaneous emission by hand:

$$i \frac{\partial C_2}{\partial t} = \Omega \cos \omega t e^{i\omega_0 t} C_1 - i\gamma_{sp} C_2$$

where $\gamma_{sp} = \frac{A_{21}}{2} = \frac{1}{2} \cdot \frac{e^2 \omega_0^3 D_{12}^2}{3\pi \epsilon_0 \hbar c^3}$

Spontaneous decay rate

natural linewidth

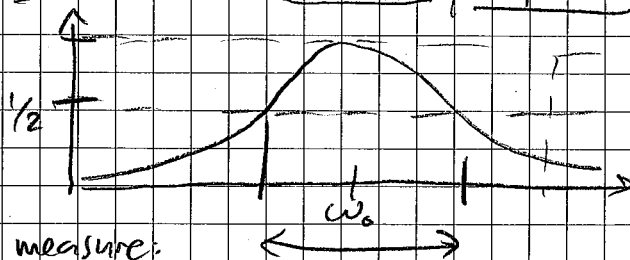
Obtain

$$C_2(t) = -\frac{1}{2} \Omega \left\{ \underbrace{\frac{e^{i(\omega_0 + \omega)t}}{\omega_0 + \omega - i\gamma_{sp}}}_{\rightarrow \text{RWA}} + \frac{e^{i(\omega_0 - \omega)t}}{\omega_0 - \omega - i\gamma_{sp}} \right\}$$

Finally obtain

$$\chi''(\omega) = \frac{2\pi N c^3}{\omega_0^3 V} \frac{\gamma_{sp}^2}{(\omega_0 - \omega)^2 + \gamma_{sp}^2}$$

a Lorentzian line profile.



Important measure:

Full width at half maximum

$$\text{FWHM} = 2\gamma_{sp}$$

4-3 Qualitative (& sloppy!) explanations:

Energy-time uncertainty relation

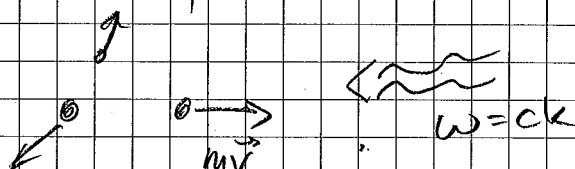
$$\Delta E \Delta t \geq \hbar$$

Δt here: $\tau_e = \frac{1}{\gamma_{sp}}$ lifetime of excited state

ΔE : linewidth $\sim$ "uncertainty in energy of exc state"

Doppler broadening

Atoms have different velocities



Light is Doppler shifted:

$$\omega_0 \rightarrow \omega = \omega_0 - v \cdot k$$

Atoms are Maxwell-Boltzmann distributed:

$$P(v_i) \propto e^{-\frac{M v_i^2 / 2}{k_B T}}$$

$\Rightarrow$ absorbed light is distributed according to

$$P(\omega) \propto e^{-\frac{M}{2k_B T} \left(c^2 \frac{(\omega - \omega_0)^2}{\omega_0^2} \right)}$$

a Gaussian distribution.

Note Fig. 2.41 Difference

Typically:

4-4

Homogeneous broadening - Same for all atoms
typically Lorentzian

Inhomogeneous broadening - Different atoms have
different resonances
Typically Gaussian

There is also collisional broadening
Power — " —

To obtain the true lineshape,
combine the effects by convolution

$$F(\omega) = \int_{-\infty}^{\infty} d\nu F_1(\nu) F_2(\omega + \omega_0 - \nu)$$

4. The Optical Bloch equations

4-5

2-7
2-9

A more convenient formulation for some QM processes is the density matrix.

Can be adapted to - many-body
statistical problems
damping

For two states, define

$$\rho_{11}(t) = |C_1(t)|^2$$

$$\rho_{22}(t) = |C_2(t)|^2$$

$$\rho_{12}(t) = C_1(t)^* C_2(t)$$

$$\rho_{21}(t) = \rho_{12}(t)^*$$

} populations

} coherences

or, if you will,

$$\rho = \begin{pmatrix} C_1 \\ C_2 \end{pmatrix} \begin{pmatrix} C_1^* & C_2^* \end{pmatrix} \quad \text{direct product}$$

$$= \begin{pmatrix} \rho_{11} & \rho_{12} \\ \rho_{21} & \rho_{22} \end{pmatrix}$$

Note: 1) $|\rho_{12}|^2 = |\rho_{21}|^2 = \rho_{12}\rho_{21} = |C_1|^2 |C_2|^2 = \rho_{11}\rho_{22}$

2) $\rho_{11} + \rho_{22} = 1$ normalization

↑
for ordinary
QM;
"pure states"

4-10 Observe puzzle:

Schrödinger picture

Degrees of freedom

Re C_1 or $|C_1|$

Im C_1 or $\arg C_1$

Re C_2 or $|C_2|$

Im C_2 or $\arg C_2$

Constraints

$$|C_1|^2 + |C_2|^2 = 1$$

4

1

= 3

Density matrix

Degrees of fr.

P_{11}

P_{22}

Re P_{12} or $|P_{12}|$

Im P_{12} or $\arg P_{12}$

Constraints

$$|P_{12}|^2 = P_{11}P_{22}$$

$$P_{11} + P_{22} = 1$$

4

2

= 2

Does that mean that we have lost information in going to density matrix?

Answer: No. Overall phase of wavefunction is redundant

$$|\psi\rangle = C_1|1\rangle + C_2|2\rangle \rightarrow e^{i\alpha}(C_1|1\rangle + C_2|2\rangle)$$

Only relative phase between C_1 and C_2 matters

This is encoded in density matrix.

4-7 Derive eqs of motion for ρ

Use

$$i \frac{dC_1}{dt} = \Omega \cos(\omega t) e^{-i\omega t} C_2$$

$$i \frac{dC_2}{dt} = \Omega \cos(\omega t) e^{i\omega t} C_1$$

and chain rule

$$\frac{dp_{11}}{dt} = \frac{d}{dt} (C_1^* C_1) = C_1^* \frac{dC_1}{dt} + C_1 \frac{dC_1^*}{dt}$$

$$= C_1^* \frac{1}{i} \Omega \cos(\omega t) e^{-i\omega t} C_2 + C_1 \frac{1}{i^*} \Omega \cos(\omega t) e^{i\omega t} C_2^*$$

$$= -i \Omega \cos(\omega t) e^{-i\omega t} p_{21} + i \Omega \cos(\omega t) e^{i\omega t} p_{12}$$

$$\boxed{\frac{dp_{11}}{dt} = i \Omega \cos(\omega t) (e^{i\omega t} p_{12} - e^{-i\omega t} p_{21})}$$

Use $p_{22} = 1 - p_{11}$

$$\Rightarrow \boxed{\frac{dp_{22}}{dt} = - \frac{dp_{11}}{dt}}$$

$$\frac{dp_{22}}{dt} = C_2^* \frac{1}{i} \Omega \cos(\omega t) e^{-i\omega t} C_2 + C_1 \frac{1}{i^*} \Omega \cos(\omega t) e^{-i\omega t} C_1^*$$

$$\boxed{\frac{dp_{22}}{dt} = i \Omega \cos(\omega t) e^{-i\omega t} (p_{11} - p_{22})}$$

$$\boxed{\frac{dp_{21}}{dt} = \left(\frac{dp_{12}}{dt} \right)^*}$$

One more step $\rightarrow$

4-8

$$\frac{dp_{11}}{dt} = i\Omega \frac{1}{2} (e^{i\omega t} + e^{-i\omega t}) (e^{i\omega t} p_{12} - e^{-i\omega t} p_{21})$$

RWA: Discard fast oscillating terms

- because we have seen they give small contrib.

discard $e^{i(\omega_0 + \omega)t}$ and $e^{-i(\omega_0 + \omega)t}$

$$\left\{ \begin{aligned} \frac{dp_{11}}{dt} &= \frac{i\Omega}{2} (e^{i(\omega_0 - \omega)t} p_{12} - e^{-i(\omega_0 - \omega)t} p_{21}) \\ \frac{dp_{22}}{dt} &= -\frac{dp_{11}}{dt} \\ \frac{dp_{12}}{dt} &= \frac{i\Omega}{2} e^{-i(\omega_0 - \omega)t} (p_{11} - p_{22}) \\ \frac{dp_{21}}{dt} &= \left(\frac{dp_{12}}{dt} \right)^* \end{aligned} \right.$$

The O.B.E.

One more simplification

Define

$$\begin{aligned} \tilde{p}_{12} &= e^{i(\omega_0 - \omega)t} p_{12} \\ \tilde{p}_{21} &= e^{-i(\omega_0 - \omega)t} p_{21} \end{aligned}$$

$$\begin{aligned} \tilde{p}_{11} &= p_{11} \\ \tilde{p}_{22} &= p_{22} \end{aligned}$$

just so notation gets prettier

Insert

$$\boxed{\frac{d\tilde{p}_{11}}{dt} = \frac{i\Omega}{2} (\tilde{p}_{12} - \tilde{p}_{21})}$$

by inspection

$$\begin{aligned}\frac{d\tilde{\rho}_{12}}{dt} &= \frac{d}{dt} \left(\rho_{12} e^{i(\omega_0 - \omega)t} \right) = \\ &= i(\omega_0 - \omega) \rho_{12} e^{i(\omega_0 - \omega)t} + e^{i(\omega_0 - \omega)t} \frac{d\rho_{12}}{dt}\end{aligned}$$

$$\left[\frac{d\tilde{\rho}_{12}}{dt} = i(\omega_0 - \omega) \tilde{\rho}_{12} + \frac{i\Omega}{2} (\tilde{\rho}_{11} - \tilde{\rho}_{22}) \right]$$

See if we can find a solution of the form

$$\tilde{\rho}_{ij}(t) = \tilde{\rho}_{ij}(0) e^{\lambda t}$$

$$\frac{i\Omega}{2} (\tilde{\rho}_{12}(0) - \tilde{\rho}_{21}(0)) = \lambda \tilde{\rho}_{11}(0)$$

$$-\frac{i\Omega}{2} (\tilde{\rho}_{12}(0) - \tilde{\rho}_{21}(0)) = \lambda \tilde{\rho}_{22}(0)$$

$$i(\omega_0 - \omega) \tilde{\rho}_{12}(0) + \frac{i\Omega}{2} (\tilde{\rho}_{11}(0) - \tilde{\rho}_{22}(0)) = \lambda \tilde{\rho}_{12}(0)$$

$$-i(\omega_0 - \omega) \tilde{\rho}_{21}(0) - \frac{i\Omega}{2} (\tilde{\rho}_{11}(0) - \tilde{\rho}_{22}(0)) = \lambda \tilde{\rho}_{21}(0)$$

an eigenvalue problem!

$$\begin{pmatrix} -\lambda & 0 & \frac{i\Omega}{2} & -\frac{i\Omega}{2} \\ 0 & -\lambda & -\frac{i\Omega}{2} & \frac{i\Omega}{2} \\ \frac{i\Omega}{2} & -\frac{i\Omega}{2} & i(\omega_0 - \omega) - \lambda & 0 \\ -\frac{i\Omega}{2} & \frac{i\Omega}{2} & 0 & -i(\omega_0 - \omega) - \lambda \end{pmatrix} \begin{pmatrix} \tilde{\rho}_{11}(0) \\ \tilde{\rho}_{22}(0) \\ \tilde{\rho}_{12}(0) \\ \tilde{\rho}_{21}(0) \end{pmatrix} = 0$$

4-10 Compute the 4x4 determinant

$$\rightarrow \begin{vmatrix} -\lambda & -\frac{i\Omega}{2} & \frac{i\Omega}{2} & 0 \\ \frac{i\Omega}{2} & i(\omega_0 - \omega) - \lambda & 0 & 0 \\ \frac{i\Omega}{2} & 0 & -i(\omega_0 - \omega) - \lambda & 0 \\ 0 & 0 & 0 & -\lambda \end{vmatrix} + \frac{1}{2} i\Omega \begin{vmatrix} 0 & -\lambda & -\frac{i\Omega}{2} \\ \frac{i\Omega}{2} & -\frac{i\Omega}{2} & i(\omega_0 - \omega) - \lambda \\ \frac{i\Omega}{2} & \frac{i\Omega}{2} & 0 \end{vmatrix} = 0$$

$$= -\lambda \left\{ -\lambda (i(\omega_0 - \omega) - \lambda) (-i(\omega_0 - \omega) - \lambda) - \left(\frac{i\Omega}{2}\right)^2 (i(\omega_0 - \omega) - \lambda) - \left(\frac{i\Omega}{2}\right)^2 (-i(\omega_0 - \omega) - \lambda) \right\}$$

$$+ \frac{1}{2} i\Omega \left\{ \left(\frac{i\Omega}{2}\right)^3 - (-\lambda) \frac{i\Omega}{2} (-i(\omega_0 - \omega) - \lambda) - \left(\frac{i\Omega}{2}\right)^3 \right\}$$

$$+ \frac{1}{2} i\Omega \left\{ -\left(\frac{i\Omega}{2}\right)^3 - \frac{i\Omega}{2} (-\lambda) (i(\omega_0 - \omega) - \lambda) - (-1) \left(\frac{i\Omega}{2}\right)^3 \right\}$$

$$= \lambda^2 \left(\lambda^2 - (i(\omega_0 - \omega))^2 \right) - \lambda^2 \cdot 2 \left(\frac{i\Omega}{2}\right)^2$$

$$+ 2 \cdot \left(\frac{i\Omega}{2}\right)^2 \lambda^2 \cdot (-1)$$

$$= \lambda^2 \left\{ \lambda^2 + (\omega_0 - \omega)^2 + 4 \frac{\Omega^2}{4} \right\}$$

$$= \lambda^2 \left\{ \lambda^2 + (\omega_0 - \omega)^2 + \Omega^2 \right\} = 0$$

Solutions: $\begin{cases} \lambda = 0 \\ \lambda = \pm i \sqrt{(\omega_0 - \omega)^2 + \Omega^2} = \pm i \Omega_{\text{gen}} \end{cases}$

"the generalized
Rabi frequency"

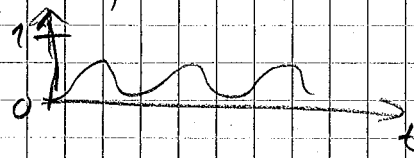
Norman

General solution

4-11

$$\rho_{ij}(t) = \rho_{ij}^{(1)} + \rho_{ij}^{(2)} e^{i\Omega_{\text{gen}} t} + \rho_{ij}^{(3)} e^{-i\Omega_{\text{gen}} t}$$

Specific solution for $\rho_{22}(0) = 0$, $\rho'_{22}(0) = 0$:

$$\rho_{22}(t) = \left(\frac{\Omega}{\Omega_{\text{gen}}}\right)^2 \sinh^2\left(\frac{1}{2}\Omega_{\text{gen}} t\right)$$


$$= \frac{1}{2} \left(\frac{\Omega}{\Omega_{\text{gen}}}\right)^2 \left(1 - \frac{1}{2} e^{i\Omega_{\text{gen}} t} - \frac{1}{2} e^{-i\Omega_{\text{gen}} t}\right)$$

$$\left(\rho_{22}^{(1)} = \frac{1}{2} \left(\frac{\Omega}{\Omega_{\text{gen}}}\right)^2 \quad \rho_{22}^{(2)} = -\frac{1}{4} \left(\frac{\Omega}{\Omega_{\text{gen}}}\right)^2 \quad \rho_{22}^{(3)} = -\frac{1}{4} \left(\frac{\Omega}{\Omega_{\text{gen}}}\right)^2\right)$$

Note: At resonance: $\Omega_{\text{gen}} = \sqrt{(\omega_0 - \omega)^2 + \Omega^2} = \Omega$

$$\text{then } \left| \begin{array}{l} \rho_{22}(t) = \sinh^2\left(\frac{1}{2}\Omega t\right) \\ \rho_{11}(t) = \cosh^2\left(\frac{1}{2}\Omega t\right) \end{array} \right|$$

"Rabi oscillations", "optical nutation"

with the Rabi frequency $\Omega = \frac{1}{\hbar} \vec{E} \cdot \vec{D}_{12}$

Ex: Let the light stay on for a time t

$$\text{s.t. } \Omega t = \pi$$

$$\text{then } \rho_{22}(t) = 1$$

$$\rho_{11}(t) = 0$$

whole population is transferred.

A "π-pulse".

Notation:

London	E.L. & rest of world
Ω	Ω
Ω	Ω_{gen}

4-12 Ex. 2. Let $\Omega T = \frac{\pi}{2}$

Then $\rho_{22}(T) = \frac{1}{2}$

$$\rho_{11}(T) = \frac{1}{2}$$

the atoms are put in a coherent superposition of $|1\rangle$ and $|2\rangle$

" $\frac{\pi}{2}$ -pulse"

useful tool for atomic physics experiments.

Spontaneous emission

Again, in the semiclassical picture have to introduce spont. em "by hand"

$$\frac{d\tilde{\rho}_{22}}{dt} = -\frac{i\Omega}{2}(\tilde{\rho}_{12} - \tilde{\rho}_{21}) - 2\gamma_{sp}\tilde{\rho}_{22} = -\frac{d\tilde{\rho}_{11}}{dt}$$

$$\frac{d\tilde{\rho}_{12}}{dt} = \frac{i\Omega}{2}(\tilde{\rho}_{11} - \tilde{\rho}_{22}) + [i(\omega_0 - \omega) - \gamma_{sp}]\tilde{\rho}_{12} = \left(\frac{d\tilde{\rho}_{21}}{dt}\right)^*$$

