

14. Laser fluctuations

What are the probabilities $P(n)$ for the laser
Steady state: $N_2 T_{st} P(n-1) = T_{cav} P(n)$ 3-level

$$\Rightarrow \dots \Rightarrow P(n) = \frac{(C n_s)^n (n_s - 1)!}{(n_s + n - 1)!} P(0)$$

Simplifies in two limits

★ Below threshold: $C < 1$

only $n < n_s$ are populated

then $(n_s + n - 1)! \approx n_s^n (n_s - 1)!$

$\Rightarrow P(n) = C^n P(0) = C^n (1 - C)$ geometric distribution

$$\text{again } \langle n \rangle = \frac{C}{1 - C}$$

$$\text{and } g^{(2)}(0) = 2$$

→ Chaotic light.

The mode is dominated by spontaneous emission.

★ Above threshold, $C > 1$

$n \gg 1$ so $n - 1 \approx n$

$$P(n) = \frac{(C n_s)^n n_s!}{(n_s + n)!} P(0)$$

$$\langle n \rangle = \sum_{n=0}^{\infty} n P(n) = \sum_n (n_s + n - n_s) \frac{(C n_s)^n n_s!}{(n_s + n)!} P(0)$$

$$= \sum_n \frac{(C n_s)^n n_s!}{(n_s + n - 1)!} P(0) - n_s \sum_n \frac{(C n_s)^n n_s!}{(n_s + n)!} P(0)$$

$$\langle n \rangle = (C n_s) \sum_{n=1}^{\infty} P(n-1) - n_s \sum_{n=1}^{\infty} P(n)$$

$$\langle n \rangle = (C-1) n_s$$

then

$$P(n) = \frac{(n_s + \langle n \rangle)^n n_s!}{(n_s + n)!} P(0)$$

$$P(n) \approx \frac{1}{\sqrt{2\pi(n_s + \langle n \rangle)}} e^{-\frac{(n - \langle n \rangle)^2}{2(n_s + \langle n \rangle)}}$$

$$\text{Variance: } (\Delta n)^2 = n_s + \langle n \rangle$$

Compare coherent state

$$P(n) = \frac{1}{\sqrt{2\pi\langle n \rangle}} e^{-\frac{(n - \langle n \rangle)^2}{2\langle n \rangle}} \quad \text{coherent}$$

This state is super-Poissonian

but close to Poissonian if $\langle n \rangle \gg n_s$

$$\text{and } g^{(2)}(0) = \frac{n_s}{\langle n \rangle^2} + 1 \approx 1$$

n_s is a little extra amplitude noise
due to spont. emission.

Rate of spontaneous emission into this mode

$$\text{is } N_2 T_{st} = \left(\begin{array}{l} \# \text{ atoms that} \\ \text{can decay} \end{array} \right) \cdot \left(\begin{array}{l} \text{decay rate of} \\ \text{each atom} \end{array} \right)$$

But we found that above threshold,

$$N_2 T_{st} = T_{\text{cav}}$$

⇒ Spontaneous photons are created at the same rate as they are lost from the cavity

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→ on average, there is one noise photon in the lasing mode.

*However, the phase will diffuse over long times

Atom-radiation dynamics (§7.7-7.9)

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Motivation: Have done rate equations.

Necessary for large- N problems - statistical method.

Now: try full quantum evolution

Suitable for problems with one atom

Work in Heisenberg picture. (4.9.24)

$$\begin{aligned}\hat{H} = & \hbar\omega_0 \hat{\pi}^\dagger(t) \hat{\pi}(t) && \text{(atom)} \\ & + \sum_{\vec{k}\lambda} \hbar\omega_k \left(\hat{a}_{\vec{k}\lambda}^\dagger(t) \hat{a}_{\vec{k}\lambda}(t) + \frac{1}{2} \right) && \text{(radiation)} \\ & + i \sum_{\vec{k}\lambda} \hbar g_{\vec{k}\lambda} \left\{ \hat{\pi}^\dagger(t) \hat{a}_{\vec{k}\lambda}(t) e^{i\vec{k}\cdot\vec{R}} - \hat{a}_{\vec{k}\lambda}^\dagger(t) \hat{\pi}(t) e^{-i\vec{k}\cdot\vec{R}} \right\} && \text{(electric dipole)}\end{aligned}$$

Put up equations of motion for the operators

$$i \frac{d\hat{a}_{\vec{k}\lambda}}{dt} = \frac{1}{\hbar} [\hat{a}_{\vec{k}\lambda}, \hat{H}] = \omega_k \hat{a}_{\vec{k}\lambda}(t) - i g_{\vec{k}\lambda} \hat{\pi}(t) e^{-i\vec{k}\cdot\vec{R}}$$

Formal solution:

$$\hat{a}_{\vec{k}\lambda}(t) = e^{-i\omega_k t} \left[\hat{a}_{\vec{k}\lambda}(0) - g_{\vec{k}\lambda} \int_0^t dt' \hat{\pi}(t') e^{-i\vec{k}\cdot\vec{R} + i\omega_k t'} \right] \quad (1)$$

$$i \frac{d\hat{\pi}}{dt} = \omega_0 \hat{\pi}(t) - i \sum_{\vec{k}\lambda} g_{\vec{k}\lambda} \left[2\hat{\pi}^\dagger(t) \hat{\pi}(t) - 1 \right] \hat{a}_{\vec{k}\lambda}(t) e^{i\vec{k}\cdot\vec{R}}$$

$$\Rightarrow \hat{\pi}(t) = e^{-i\omega_0 t} \left[\hat{\pi}(0) - \sum_{\vec{k}\lambda} g_{\vec{k}\lambda} \int_0^t dt' \left(2\hat{\pi}^\dagger(t') \hat{\pi}(t') - 1 \right) \hat{a}_{\vec{k}\lambda}(t') \times e^{i\vec{k}\cdot\vec{R} + i\omega_0 t'} \right]$$

↑
free atom
↑
coupling to field

How do we obtain a solution on closed form? (14-5)

- Proceed by iteration.

Assume $g_{\mathbf{k}\lambda}$ small.

Example: Calculate the time dependence of

$$\langle \hat{\pi}^\dagger \hat{\pi} \rangle = \langle 4|2 \rangle \langle 2|4 \rangle = |\langle 2|4 \rangle|^2$$

= degree of excitation.

$$\frac{d}{dt} (\hat{\pi}^\dagger(t) \hat{\pi}(t)) = \sum_{\mathbf{k}\lambda} g_{\mathbf{k}\lambda} \left[\hat{\pi}^\dagger(t) \hat{a}_{\mathbf{k}\lambda}(t) e^{i\vec{k} \cdot \vec{R}} + \hat{a}_{\mathbf{k}\lambda}^\dagger \hat{\pi}(t) e^{-i\vec{k} \cdot \vec{R}} \right] \quad (3)$$

[again, this can be found by working out the commutator]

Insert $\hat{\pi}^\dagger, \hat{\pi}, \hat{a}^\dagger, \hat{a}$ to lowest order in $g_{\mathbf{k}\lambda}$

$$(2) \Rightarrow \hat{\pi}(t) = e^{-i\omega_0 t} \hat{\pi}(0) \Rightarrow \hat{\pi}(t') = e^{-i\omega_0(t'-t)} \hat{\pi}(t)$$

into (1)

$$(1) \Rightarrow \hat{a}_{\mathbf{k}\lambda}(t) = e^{-i\omega_k t} \left(\hat{a}_{\mathbf{k}\lambda}(0) - g_{\mathbf{k}\lambda} \hat{\pi}(t) e^{i\omega_0 t} e^{-i\vec{k} \cdot \vec{R}} \int_0^t dt' e^{i(\omega_k - \omega_0)t'} \right)$$

$$(3) \Rightarrow \frac{d}{dt} (\hat{\pi}^\dagger \hat{\pi}) = \hat{\pi}^\dagger(t) \sum_{\mathbf{k}\lambda} g_{\mathbf{k}\lambda} \left[\hat{a}_{\mathbf{k}\lambda}(0) e^{i\vec{k} \cdot \vec{R} - i\omega_k t} - g_{\mathbf{k}\lambda} \hat{\pi}(t) \int_0^t dt' e^{i(\omega_k - \omega_0)(t'-t)} \right] + \text{h.c.}$$

Now take the expectation value in vacuum for the radiation:

$|0_{\text{rad}}\rangle$, (no photons)

$$\langle 0_{\text{rad}} | \hat{a}_{\mathbf{k}\lambda}(t) | 0_{\text{rad}} \rangle = 0, \text{ so}$$

$$\frac{d}{dt} \langle 0_{\text{rad}} | \hat{\pi}^\dagger \hat{\pi} | 0_{\text{rad}} \rangle = - \langle 0_{\text{rad}} | \hat{\pi}^\dagger(t) \hat{\pi}(t) | 0_{\text{rad}} \rangle \sum_{\mathbf{k}\lambda} g_{\mathbf{k}\lambda}^2 \underbrace{2 \frac{\sin[(\omega_k - \omega_0)t]}{\omega_k - \omega_0}}_{\rightarrow \pi \delta(\omega_k - \omega_0)}$$

as $t \rightarrow \infty$

$$\therefore \frac{d}{dt} \underbrace{\langle 0_r | \hat{\pi}^\dagger \hat{\pi} | 0_r \rangle}_{= \rho_{22}} = - \underbrace{\langle 0_r | \hat{\pi}^\dagger(t) \hat{\pi}(t) | 0_r \rangle \cdot 2\pi \sum_{k\lambda} g_{k\lambda}^2 \delta(\omega_k - \omega_0)}_{= 2\gamma_{sp}} \quad \boxed{14-6}$$

$\therefore$ We have reproduced the result

$$\frac{d\rho_{22}}{dt} = -2\gamma_{sp}\rho_{22} \quad \text{in vacuum.}$$

• What about the rate of atomic transition?

$$\begin{aligned} \frac{d}{dt} \langle 0_r | \hat{\pi}(t) | 0_r \rangle &= - \underbrace{\langle 0_r | \hat{\pi}(t) | 0_r \rangle \left\{ i\omega_0 + \sum_{k\lambda} g_{k\lambda}^2 \int_0^t dt' e^{i(\omega_k - \omega_0)(t'-t)} \right\}}_{i\omega_0 + \sum_{k\lambda} g_{k\lambda}^2 \left(\frac{1 - \cos[(\omega_k - \omega_0)t]}{i(\omega_k - \omega_0)} + \frac{\sin(\omega_k - \omega_0)t}{\omega_k - \omega_0} \right)} \\ &= i\omega_0 - i \underbrace{\left[\sum_{\lambda} \frac{V}{(2\pi)^3} \int d\vec{k} \lim_{t \rightarrow \infty} \left(\frac{1 - \cos[(\omega_k - \omega_0)t]}{i(\omega_k - \omega_0)} \right) \cdot g_{i\lambda}^2 \right]}_{\text{vacuum correction}} + \gamma_{sp} \end{aligned}$$

A vacuum correction to the
resonance frequency
"the Lamb shift"

Triumph of QED: this small shift $\sim 10^{-6}$
measured by Lamb in 1950s

★ What is $\hat{\vec{E}}_T^+(\vec{r}, t)$ in the same approx.?

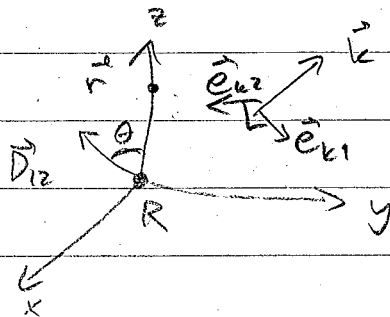
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$$\hat{\vec{E}}_T^+(\vec{r}, t) = i \sum_{\vec{k}, \lambda} \vec{e}_{\vec{k}, \lambda} \sqrt{\frac{\hbar \omega_k}{2 \epsilon_0 V}} \hat{a}_{\vec{k}, \lambda}(t) e^{i \vec{k} \cdot \vec{r}}$$

Insert our ^{exact} expression for $\hat{a}_{\vec{k}, \lambda}(t)$
neglect the $\hat{a}(0)$ term

$$\hat{\vec{E}}_T^+(\vec{r}, t) = -i \sum_{\vec{k}, \lambda} \vec{e}_{\vec{k}, \lambda} \underbrace{\sqrt{\frac{\hbar \omega_k}{2 \epsilon_0 V}} g_{\vec{k}, \lambda}}_{\frac{e}{16 \pi^3 \epsilon_0} \omega_k \vec{e}_{\vec{k}, \lambda} \cdot \vec{D}_{12}} e^{i \vec{k} \cdot (\vec{r} - \vec{R})} \int_0^t dt' \hat{\pi}(t') e^{i \omega_k t'}$$

Note:



$\vec{r}$ is the point of observation

$\vec{R}$ is the location of the nucleus

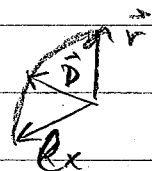
$$\vec{k} \perp \vec{e}_{k1} \perp \vec{e}_{k2}$$

this is called the source-field expression $\vec{E}_{sf}$ because it retains only the part coming from the source (=atom), containing $\hat{\pi}$ not the free-field part $\hat{a}(0)$

Some more manipulation yields

$$\hat{\vec{E}}_{sf}(\vec{r}, t) = - \frac{e \omega_0^2 D_{12} \sin \Theta}{4 \pi \epsilon_0 c^2 |\vec{r} - \vec{R}|} \vec{e}_x \hat{\pi} \left(t - \frac{|\vec{r} - \vec{R}|}{c} \right)$$

where $\vec{e}_x$ is a unit vector for the projection of $\vec{D}_{12}$ perpendicular to $\vec{r} - \vec{R}$.



Emission by a single driven atom

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Assume the light is polarized onto $\vec{e}_x$:

$$\hat{\vec{E}}_{sf} \rightarrow \hat{E}_{sf} \text{ scalar.}$$

Coherences

$$g^{(1)}(\tau) = \frac{\langle E_{sf}^-(\vec{r}, t) E_{sf}^+(\vec{r}, t+\tau) \rangle}{\langle E_{sf}^-(\vec{r}, t) E_{sf}^+(\vec{r}, t) \rangle} = \frac{\langle \hat{\pi}^+(t) \hat{\pi}(t+\tau) \rangle}{\langle \hat{\pi}^+(t) \hat{\pi}(t) \rangle}$$

$$g^{(2)}(\tau) = \frac{\langle \hat{\pi}^+(t) \hat{\pi}^+(t+\tau) \hat{\pi}(t+\tau) \hat{\pi}(t) \rangle}{\langle \hat{\pi}^+(t) \hat{\pi}(t) \rangle^2}$$

For zero delay $\tau=0$

$$g^{(1)}(0) = 1$$

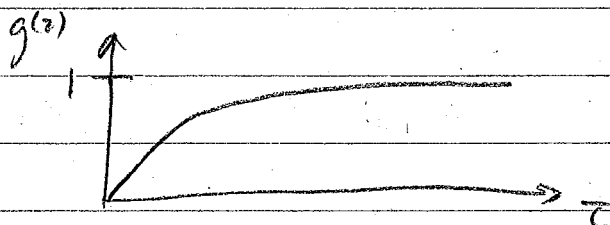
$$g^{(2)}(0) = 0 \quad \text{because } \hat{\pi}(t) \hat{\pi}(t) = 0$$

Sub-Poissonian!

How calculate $\langle \hat{\pi}^+(t) \hat{\pi}^+(t+\tau) \hat{\pi}(t+\tau) \hat{\pi}(t) \rangle$

Assume: After a long enough time delay τ ,
 $\hat{\pi}^+(t+\tau) \hat{\pi}(t+\tau)$ and $\hat{\pi}^+(t) \hat{\pi}(t)$ become
independent operators, and

$$g^{(2)}(\tau) \rightarrow \frac{\langle \hat{\pi}^+(t) \hat{\pi}(t) \rangle \langle \hat{\pi}^+(t+\tau) \hat{\pi}(t+\tau) \rangle}{\langle \hat{\pi}^+(t) \hat{\pi}^+(t) \rangle^2} = 1 \quad \text{for } \tau \rightarrow \infty$$



Anti-bunched.

London applies a mixture of OBE
and other things to arrive at

$$g^{(2)}(\tau) = 1 - e^{-2\gamma_{sp}|\tau|}$$

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