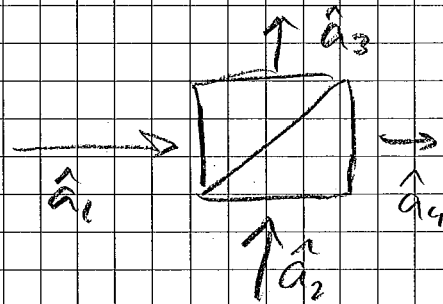


10. Quantum-mechanical beam-splitter theory



Virtue of beam splitter: create superposition states!

Reflection coeff. R

Transmission T

$$\begin{cases} |R|^2 + |T|^2 = 1 \\ RT^* + TR^* = 0 \end{cases}$$

(derived in Ch. 3)

$$\hat{a}_3 = R\hat{a}_1 + T\hat{a}_2$$

$$\hat{a}_4 = T\hat{a}_1 + R\hat{a}_2$$

Reverse:

$$\hat{a}_1 = R^*\hat{a}_3 + T^*\hat{a}_4$$

$$\hat{a}_2 = T^*\hat{a}_3 + R^*\hat{a}_4$$

Independent input fields: $[\hat{a}_1, \hat{a}_2^\dagger] = 0$

The electric fields

$$\hat{E}_3(x) = |R|\hat{E}_1(x - \phi_R) + |T|\hat{E}_2(x - \phi_T)$$

$$\hat{E}_4(x) = |T|\hat{E}_1(x - \phi_T) + |R|\hat{E}_2(x - \phi_R)$$

where ϕ_R and ϕ_T are the phases of the complex numbers R and T

10-2

Photon numbers

$$\hat{n}_3 = |R|^2 \hat{n}_1 + R^* T \hat{a}_1^\dagger \hat{a}_2 + R T^* \hat{a}_2^\dagger \hat{a}_1 + |T|^2 \hat{n}_2$$

$$\hat{n}_4 = |T|^2 \hat{n}_1 + R T^* \hat{a}_1^\dagger \hat{a}_2 + R^* T \hat{a}_2^\dagger \hat{a}_1 + |R|^2 \hat{n}_2$$

& variances (assuming arm 2 in vacuum state, for simplicity)

$$(\Delta n_3)^2 = |R|^4 (\Delta n_1)^2 + |R|^2 |T|^2 \langle n_1 \rangle$$

$$(\Delta n_4)^2 = |T|^4 (\Delta n_1)^2 + |T|^2 |R|^2 \langle n_1 \rangle$$

• Single-photon input:

One photon in arm 1, vacuum in arm 2:

$$|1\rangle_1 |0\rangle_2 = \hat{a}_1^\dagger |0\rangle$$

But $\hat{a}_1^\dagger = R \hat{a}_3^\dagger + T \hat{a}_4^\dagger$

$$|1\rangle_1 |0\rangle_2 = R |1\rangle_3 |0\rangle_4 + T |0\rangle_3 |1\rangle_4$$

the BS creates an entangled state.

(a very important concept in quantum information)

Recall quantum measurement postulate:

A measurement collapses the state onto the corresponding eigenstate



If we look in arm 3 and see no photons, then

$$|A_{\text{meas}}\rangle = N \cdot {}_3\langle 0 | \{ R |1\rangle_3 |0\rangle_4 + T |0\rangle_3 |1\rangle_4 \} = |1\rangle_4$$

Expectation values of photon number

10-3

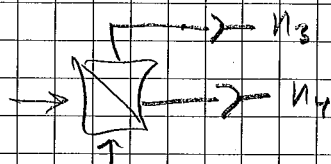
corresponds to average over many measurements:

$$\begin{aligned}\langle n_3 \rangle &= \langle 0_2 | \langle 1_1 | \hat{n}_3 | 1_1 \rangle | 0_2 \rangle \\ &= \langle 0_2 | \langle 1_1 | (R^\dagger a_1^\dagger + T^\dagger a_2^\dagger) (R \hat{a}_1 + T \hat{a}_2) | 1_1 \rangle | 0_2 \rangle \\ &= |R|^2\end{aligned}$$

$$\langle n_4 \rangle = |T|^2$$

that was not so hard to guess, was it?

• Hanbury Brown-Twiss experiment



measures $\langle n_3 n_4 \rangle$

For one-photon input, get

$$\langle n_3 n_4 \rangle = \dots = 0$$

... could never happen in classical theory!

Recall classical:

$$\langle I_3 \rangle = |R|^2$$

$$\langle I_4 \rangle = |T|^2$$

$$\langle I_3(t) I_4(t+\tau) \rangle = \langle I_3 \rangle \langle I_4 \rangle g^{(2)}(\tau)$$

$$\langle I_3(t) I_4(t) \rangle = \langle I_3 \rangle \langle I_4 \rangle \underbrace{g^{(2)}(0)}$$

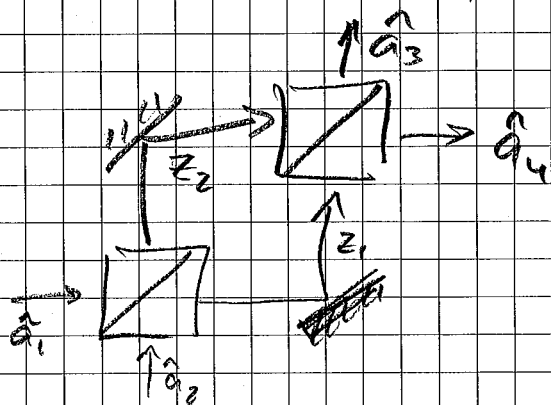
= 1 for stable wave

= 2 for chaotic light

→ arbitrary input

10-4

Mach-Zehnder interferometer



$$\hat{a}_3 = R_{M2} \hat{a}_1 + T_{M2} \hat{a}_2 \quad \hat{a}_4 = T_{M2} \hat{a}_1 + R'_{M2} \hat{a}_2$$

$$\begin{cases} R_{M2} = R^2 e^{ikz_1} + T^2 e^{ikz_2} & R'_{M2} = T^2 e^{ikz_1} + R^2 e^{ikz_2} \\ T_{M2} = RT(e^{ikz_1} + e^{ikz_2}) \end{cases}$$

Calculate mean photon number in arm 4

$$\langle n_4 \rangle = |T_{M2}|^2 = |R|^2 |T|^2 \cdot 4 \cos^2[k(z_1 - z_2)/2]$$

= same as classical! interference fringes.

• For arbitrary input $|n_1\rangle$ in arm 1,
obtain

$$\langle n_4 \rangle = |R|^2 |T|^2 \langle n \rangle 4 \cos^2[k(z_1 - z_2)/2]$$

$\Rightarrow$ it does not matter if we have many photons simultaneously, or one at a time, several times.

$\therefore$ interference is a single-photon effect!

Each single photon excites both paths of the MZ.

It does not go either way.

• Now, let arm 1 be in an arbitrary state

(10-5)

& arm 2 still in vacuum $|n_1\rangle|0_2\rangle$

We get for single BS:

$$\langle n_3 \rangle = \langle 0_2 | \langle n_1 | (|R|^2 a^\dagger a_1) | n_1 \rangle | 0_2 \rangle$$

$$= |R|^2 \langle n_1 \rangle$$

$$\langle n_4 \rangle = |T|^2 \langle n_1 \rangle$$

HBT experiment:

$$\langle n_3 n_4 \rangle = \dots = |R|^2 |T|^2 \langle n_1 (n_1 - 1) \rangle$$

We can define a correlation function

$$g_{ab}^{(2)}(\tau) = \frac{\langle \hat{a}_a^\dagger \hat{a}_a^\dagger \hat{a}_b \hat{a}_b \rangle}{\langle \hat{a}_a^\dagger \hat{a}_a \rangle \langle \hat{a}_b^\dagger \hat{a}_b \rangle}$$

Remember
this was
time-independent
in single-mode case!

then

$$g_{34}^{(2)} = \frac{\langle n_3 n_4 \rangle}{\langle n_3 \rangle \langle n_4 \rangle}$$

and

$$g_{11}^{(2)} = \frac{\langle a^\dagger a^\dagger a a \rangle}{\langle n_1 \rangle^2} = \frac{\langle n_1 (n_1 - 1) \rangle}{\langle n_1 \rangle^2}$$

use $[a, a^\dagger] = 1$

Observe

$$g_{34}^{(2)}(\tau) = \frac{\langle n_3 n_4 \rangle}{\langle n_3 \rangle \langle n_4 \rangle} = \frac{|R|^2 |T|^2 \langle n_1 (n_1 - 1) \rangle}{|R|^2 \langle n_1 \rangle |T|^2 \langle n_1 \rangle} = g_{11}^{(2)}(\tau)!$$

classical result.

10.6

Coherent-state input:

$$\text{Let } |4_i\rangle = \hat{D}_4(\alpha) |0\rangle = e^{\alpha \hat{a}_1^\dagger + \alpha^* \hat{a}_1} |0\rangle$$

$$= \exp\{\alpha(R\hat{a}_3^\dagger + T\hat{a}_4^\dagger) + \alpha^*(R^*\hat{a}_3 + T^*\hat{a}_4)\} |0\rangle$$

$$= \hat{D}_3(R\alpha) \hat{D}_4(T\alpha) |0\rangle$$

Observe!
$$e^{\alpha R\hat{a}_3^\dagger + \alpha^* R^*\hat{a}_3 + \alpha T\hat{a}_4^\dagger + \alpha^* T^*\hat{a}_4}$$

$$= e^{\alpha R\hat{a}_3^\dagger + \alpha^* R^*\hat{a}_3} \cdot e^{\alpha T\hat{a}_4^\dagger + \alpha^* T^*\hat{a}_4}$$

We could factorize only because 3 & 4 are independent: $[\hat{a}_3, \hat{a}_4^\dagger] = 0$

$$|4_i\rangle = |R\alpha_3\rangle |T\alpha_4\rangle$$

then

$$\langle n_3 n_4 \rangle = |R\alpha|^2 |T\alpha|^2 = \langle n_3 \rangle \langle n_4 \rangle$$

but $\langle n_3 n_4 \rangle = |R|^2 |T|^2 \langle n_i (n_i - 1) \rangle$

$$= |R|^2 |T|^2 (\langle n_i^2 \rangle - \langle n_i \rangle)$$

$$= |R|^2 |T|^2 \langle n_i \rangle^2$$

$$\left\{ \begin{array}{l} \langle n_i^2 \rangle = \langle n_i \rangle^2 + \langle n_i \rangle \\ \text{for coh. states!} \end{array} \right\}$$

$\therefore$ Same result as 2 independent sources.

= same as coherent classical light.

Nonclassical light

10.7

Poissonian distribution: $\langle n^2 \rangle = \langle n \rangle$

A prob. distribution is called super-Poissonian if $\langle n^2 \rangle \geq \langle n \rangle$

and sub-Poissonian if $\langle n^2 \rangle \leq \langle n \rangle$

Classical light is always super-P.

$$1 \leq g^{(2)}(0) \leq \infty \quad (\text{classical})$$

Quantum:

$$1 - \frac{1}{\langle n \rangle} \leq g^{(2)}(0) < \infty \quad \text{for } \langle n \rangle \geq 1$$

$$0 \leq g^{(2)}(0) \quad \text{for } \langle n \rangle < 1 \text{ \& \textit{single-mode.}}$$

[Disregard Eq. (5.10.4), it's just confusing]

Classically: $\langle I^2 \rangle$

May measure I & leave it undisturbed

→ measure it again

Quantum: $\langle n(n-1) \rangle$

First measurement absorbs a photon

* ... so q.m. light can be sub-Poissonian

→ e.g. in photon number states.

